# **Inorganic Chemistry**

# Efficient Near-Infrared Downconversion and Energy Transfer Mechanism of Ce<sup>3+</sup>/Yb<sup>3+</sup> Codoped Calcium Scandate Phosphor

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**ABSTRACT:** An efficient near-infrared (NIR) downconversion has been demonstrated in  $CaSc_2O_4$ :  $Ce^{3+}/Yb^{3+}$  phosphor. Doping concentration optimized  $CaSc_2O_4$ :  $1\%Ce^{3+}/5\%Yb^{3+}$  shows stronger NIR emission than doping concentration also optimized typical YAG:  $1\%Ce^{3+}/5\%Yb^{3+}$  under 470 nm excitation. The NIR emission from 900 to 1100 nm is enhanced by a factor of 2.4. In addition, the main emission peak of Yb<sup>3+</sup> in the  $CaSc_2O_4$  around 976 nm matches better with the optimal spectral response of the c-Si solar cell. The visible and NIR spectra and the decay curves of  $Ce^{3+}$ :  $5d \rightarrow 4f$  emission were used to demonstrate the energy transfer from  $Ce^{3+}$  ions to Yb<sup>3+</sup> ions. The downconversion phenomenon has been observed under the direct excitation of  $Ce^{3+}$  ions. On analyzing the dependence of energy transfer rate on Yb<sup>3+</sup> ion concentration, we reveal that the energy transfer (ET) from  $Ce^{3+}$  ions to Yb<sup>3+</sup> ions in  $CaSc_2O_4$  occurs mainly by the single-step ET process. Considering that the luminescence efficiency of  $CaSc_2O_4$ :  $Ce^{3+}$  is comparable to that of



commercial phosphor YAG:  $Ce^{3+}$ , the estimated maximum energy transfer efficiency reaches 58% in the  $CaSc_2O_4$ :  $1\%Ce^{3+}/15\%$  Yb<sup>3+</sup> sample, indicating that  $CaSc_2O_4$ :  $Ce^{3+}/Yb^{3+}$  sample has the potential in improving the conversion efficiency of c-Si solar cells.

# ■ INTRODUCTION

Due to converting sunlight into electricity directly, photovoltaic solar cells are considered as the prime candidates for effective large-scale capture and conversion of the sustainable solar energy.<sup>1,2</sup> At present, the commercial crystalline silicon solar cell (c-Si) dominating the photovoltaic market has an efficiency of only 18%, because the terrestrial spectrum of the solar radiation on the surface of the earth (AM 1.5 G) has a large energy mismatch with the response spectrum of silicon solar cells.<sup>2,3</sup> Silicon solar cells have the strongest absorption around 800-1100 nm.<sup>4</sup> Due to higher reflection and absorption, the response becomes weak in the short wavelength region, especially in the 300-500 nm region.<sup>5</sup> However, the solar spectrum is still very strong in this region.<sup>6</sup> Downconversion (DC) materials have provided a promising way for the enhancement of solar cell efficiency, which can convert the high-energy photons to near-infrared (NIR) photons.<sup>7</sup> Choosing the suitable DC material, the actual efficiency of a silicon solar cell can be improved to about 40%,<sup>7</sup> well over the efficiency limit (29%) estimated by Shockley and Queisser.<sup>8</sup> DC has also been reported in several RE<sup>3+</sup>/Yb<sup>3+</sup> couples, such as  $Pr^{3+}/Yb^{3+}$ ,  $Tm^{3+}/Yb^{3+}$ ,  $Tb^{3+}/Yb^{3+}$ ,  $Er^{3+}/Yb^{3+}$ ,  $Nd^{3+}/Yb^{3+}$ ,  $Ho^{3+}/Yb^{3+}$ , and so on.<sup>9-14</sup> Yb<sup>3+</sup> ion is suitable as an acceptor and emitter because it has an excited state at approximately 10 000 cm<sup>-1</sup>, just above the band edge of crystalline silicon.<sup>15</sup> However, the absorption transition of Pr<sup>3+</sup>, Tb<sup>3+</sup>, Er<sup>3+</sup>, Ho<sup>3+</sup>, and  $Tm^{3+}$  as a sensitizer arises from the forbidden  $4f \rightarrow 4f$  transitions, which are naturally weak in intensity and narrow in bandwidth.<sup>6,16</sup> Unlike these RE ions, the absorption of Ce<sup>3+</sup> is originated from the allowed 4f  $\rightarrow$  5d electric-dipole transitions in the broad spectral range, which results in a very high absorption cross section of about  $10^{-18}~{\rm cm}^2.^{16}$  The value of the 4f  $\rightarrow$  4f transitions of RE<sup>3+</sup> donor ions is merely  $10^{-21}~{\rm cm}^2.^{16}$  Moreover, the excitation and emission spectra of Ce<sup>3+</sup> can be tuned by the host materials, which means that DC from Ce<sup>3+</sup> to Yb<sup>3+</sup> can be enhanced by choosing appropriate host materials.<sup>17</sup> The DC of Ce<sup>3+</sup>/Yb<sup>3+</sup> combination has been reported in various oxide hosts, such as Y\_3Al\_5O\_{12}, borates, silicates, and phosphates.<sup>6,18–21</sup>

 $CaSc_2O_4$  as the host lattice which has an orthorhombic  $CaFe_2O_4$  structure with space group *Pnam* (No. 62) exhibits low crystal field symmetry and strong crystal field intensity.<sup>22</sup> Low-symmetry hosts exert a crystal field containing more uneven components around the dopant ions which can improve the transition probabilities. The strong crystal field can lead to the large Stark splitting of the energy level of dopant ions, which can improve the energy transfer efficiency. Furthermore,  $CaSc_2O_4$  has a low cutoff phonon frequency of 540 cm<sup>-1</sup>. The low cutoff phonon frequency can inhibit the nonradiative multiphonon relaxation in the DC process.<sup>22</sup> It has been reported as a promising host for achieving efficient DC

Received: February 4, 2015 Published: May 6, 2015 luminescence in some DC materials such as the green-emitting  $CaSc_2O_4$ :  $Ce^{3+}$  and  $CaSc_2O_4$ :  $Tb^{3+}$  and red-emitting  $CaSc_2O_4$ :  $Eu^{3+,23-25}$   $CaSc_2O_4$  host is proper for doped  $Ce^{3+}$  ions, in which the energy of the  $4f \rightarrow 5d$  electric-dipole transitions for  $Ce^{3+}$  locates in the visible region. Besides, the absorption region also matches well with the high-energy part of the solar spectrum, where the solar cells conversion efficiency is very low. In particular, for  $CaSc_2O_4$ :  $Ce^{3+}$  phosphor, its intense green emission under blue light excitation is comparable to the commercial  $Y_3Al_5O_{12}$  (YAG):  $Ce^{3+}$  phosphor.<sup>23</sup> The  $Ce^{3+}/Yb^{3+}$  codoped YAG material as an efficient conversion material for solar cell applications has been studied recently.<sup>6,18</sup> To our knowledge, the DC properties of  $Ce^{3+}/Yb^{3+}$  in  $CaSc_2O_4$  have not been demonstrated yet.

In the current work, we report the efficient NIR DC in doping concentration optimized  $CaSc_2O_4$ :  $1\%Ce^{3+}/5\%Yb^{3+}$ sample, and the NIR emission from 900 to 1100 nm is enhanced by a factor of 2.4, in comparison with that in doping concentration also optimized typical phosphor YAG:  $1\%Ce^{3+}/$  $5\%Yb^{3+}$ . We study the effect of Yb<sup>3+</sup> concentration on emission spectra, decay lifetime, energy transfer efficiency, and energy transfer mechanism in  $CaSc_2O_4$ :  $Ce^{3+}/Yb^{3+}$  sample. Results show  $CaSc_2O_4$ :  $Ce^{3+}/Yb^{3+}$  phosphor is a proper material for application in solar cells.

### EXPERIMENTAL SECTION

**Sample Preparation.** The  $CaSc_2O_4$ :  $1\%Ce^{3+}/x\%Yb^{3+}$  powder samples were synthesized by a solid state reaction.<sup>25</sup> The constituent oxides or carbonates  $CaCO_3$ ,  $Sc_2O_3$ ,  $Yb_2O_3$ , and  $CeO_2$  were employed as raw materials, which were mixed homogeneously by an agate mortar for 30 min and placed in a crucible with a lid. The crucible was buried by carbon sticks and sintered at 1500 °C for 4 h in CO-reducing atmosphere.

For comparison, in parallel, the YAG:  $1\%Ce^{3+}/x\%Yb^{3+}$  powder samples were also prepared by the conventional solid state reaction method.  $Y_2O_3$ ,  $Al_2O_3$ ,  $Yb_2O_3$ , and  $CeO_2$  were mixed in a ratio molar of  $(Y_{1-1\%-x\%}Ce_{1\%}Yb_{x\%})_3Al_5O_{12}$ . After a good mixing in an agate mortar, the mixture was sintered at 1600 °C for 5 h under a CO reducing atmosphere. Both  $CaSc_2O_4$  and YAG samples produced exhibit high crystallinity as evidenced by XRD data. The two powder samples  $CaSc_2O_4$ :  $1\%Ce^{3+}/5\%Yb^{3+}$  and YAG:  $1\%Ce^{3+}/5\%Yb^{3+}$  have been optimized to obtain the highest NIR emission intensity of Yb<sup>3+</sup> in each host. It turns out that the optimal fractional concentrations of the dopants in the two hosts are the same.

**Measurements and Characterization.** Powder X-ray diffraction (XRD) data was collected using Cu K $\alpha$  radiation ( $\lambda = 1.54056$  Å) on a Bruker D8 advance diffractometer. The excitation and emission spectra were measured using a FLS920 spectrometer (Edinburgh Instruments, U.K.). The fluorescence decay curves of Ce<sup>3+</sup> were measured by a FL920 fluorescence lifetime spectrometer (Edinburgh Instruments Ltd.). All of the measurements above were performed at room temperature.

#### RESULTS AND DISCUSSION

The crystal structures characterized by the XRD patterns are shown in Figure 1a for phosphors with nominal compositions of CaSc<sub>2</sub>O<sub>4</sub>: 1%Ce<sup>3+</sup>/x%Yb<sup>3+</sup> (x = 0, 5, 15, 30). All experimental XRD patterns of the samples are indexed well to the standard cards of JCPD card 20-0234. CaSc<sub>2</sub>O<sub>4</sub> has an orthorhombic CaFe<sub>2</sub>O<sub>4</sub> structure with space group *Pnam* (No. 62) and lattice constants  $a \neq b \neq c$ ,  $\alpha = \beta = \gamma = 90^{\circ}$ , as shown in Figure 1b. In this structure, Ca<sup>2+</sup> is 8-fold with a large effective ionic radius (1.12 Å) and Sc<sup>3+</sup> 6-fold with a small radius (0.745 Å). The unit cell of CaSc<sub>2</sub>O<sub>4</sub> shows there is only one site for Ca<sup>2+</sup> and two coordinated sites for Sc<sup>3+</sup>. The



**Figure 1.** (a) XRD patterns for samples with nominal compositions of  $CaSc_2O_4$ :  $1\%Ce^{3+}/x\%Yb^{3+}$  (*x* = 0, 5, 15, 30). (b) Unit cell of  $CaSc_2O_4$ .

occupied site of Ce<sup>3+</sup> is a Ca<sup>2+</sup> site in the CaSc<sub>2</sub>O<sub>4</sub>; the ionic radius of Ce<sup>3+</sup> (1.143 Å) is close to that of Ca<sup>2+,23</sup> The Yb<sup>3+</sup> ions with a large ionic radius (0.868 Å) occupy Sc<sup>3+</sup> sites in CaSc<sub>2</sub>O<sub>4</sub>, which has been reported in Yb<sup>3+</sup> singly doped CaSc<sub>2</sub>O<sub>4</sub> sample.<sup>26</sup> No impurity phase is detected in all XRD patterns, demonstrating that Ca<sup>3+</sup> and Yb<sup>3+</sup> substitutions have no effect on the phase structure of CaSc<sub>2</sub>O<sub>4</sub> host.

In order to investigate the DC properties of  $Ce^{3+}/Yb^{3+}$ codoped  $CaSc_2O_4$  phosphor, the excitation (PLE) and emission (PL) spectra of samples as direct evidence of  $Ce^{3+} \rightarrow Yb^{3+}$ energy transfer (ET) were measured and are shown in Figures 2 and 3, respectively. In Figure 2, the intense excitation band centered at around 466 nm for monitoring  $Ce^{3+}$  emission at



Figure 2. Excitation spectra of  $Ce^{3+}$  emission centered at 511 nm (black line) and Yb<sup>3+</sup> emission at 1004 nm (red line) in  $CaSc_2O_4$ : 1%  $Ce^{3+}/5\%$ Yb<sup>3+</sup> sample.

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**Figure 3.** Visible and NIR emission spectra of the samples with nominal compositions of  $CaSc_2O_4$ :  $1\%Ce^{3+}/x\%Yb^{3+}$  (x = 0, 1, 2, 5, 10, 15, 30) under 466 nm excitation.

511 nm can be assigned to the broad-band absorption of Ce<sup>3+</sup>:  $4f \rightarrow 5d$ . The Ce<sup>3+</sup>:  $4f \rightarrow 5d$  absorption is also observed by monitoring Yb<sup>3+</sup> emission at 1004 nm, which indicates the presence of Ce<sup>3+</sup>  $\rightarrow$  Yb<sup>3+</sup> ET.

The emission spectra covering the visible and near-infrared spectral range for samples with a fixed  $Ce^{3+}$  concentration (1%) and various concentrations of Yb<sup>3+</sup> are presented under 466 nm excitation corresponding to  $Ce^{3+}$ : 4f  $\rightarrow$  5d absorption in Figure 3. In the visible region (450-700 nm), the broad emission band observed around 511 nm originates from the allowed electronic transition of the excited state 5d to ground state 4f of  $\mathrm{Ce}^{3+}$  ions. The  $\mathrm{Ce}^{3+}$  emission can be decomposed into two Gaussian bands centered at about 511 nm (19 570 cm<sup>-1</sup>) and 563 nm (17762  $\text{cm}^{-1}$ ), which are attributed to the transitions from the lowest Stark level of the 5d excited state to different Stark levels  $({}^{2}F_{5/2}$  and  ${}^{2}F_{7/2})$  of the 4f ground state of Ce<sup>3+</sup>, respectively. The energy difference is about 1808 cm<sup>-1</sup>, which is very close to the splitting of the 4f ground state of Ce<sup>3+</sup> ion.<sup>2</sup> In the NIR region (900-1100 nm), a strong emission band around 1000 nm is observed for the Ce<sup>3+</sup> and Yb<sup>3+</sup> codoped samples, which is originated from the Yb<sup>3+</sup>:  ${}^{2}F_{5/2} \rightarrow {}^{2}F_{7/2}$ transition. The appearance of the NIR emission of Yb<sup>3+</sup> under direct excitation of  $Ce^{3+}$  is another evidence for the existence of ET from  $Ce^{3+}$  to  $Yb^{3+}$ . It is noticed that the emission intensity of Ce<sup>3+</sup>: 5d  $\rightarrow$  4f has a decline with increased Yb<sup>3+</sup> ion concentration. Simultaneously, the NIR emission of  $Yb^{3+}$  first increases and then decreases when the Yb3+ concentration exceeds 5% due to concentration quenching.<sup>11</sup> The decreased emission intensity of Ce<sup>3+</sup> and increased NIR emission intensity of Yb<sup>3+</sup> when Yb<sup>3+</sup> concentration increases from 0% to 5% are also proof for the existence of ET from  $Ce^{3+}$  to  $Yb^{3+}$ . In  $Ce^{3+}/$ Yb<sup>3+</sup> codoped CaSc<sub>2</sub>O<sub>4</sub> material, x = 5% represents optimal Yb<sup>3+</sup> doped concentration which is the same as the reported typical phosphor YAG: Ce<sup>3+</sup>/Yb<sup>3+</sup>.<sup>6</sup>

The PLE and PL spectra confirm that ET from  $Ce^{3+}$  to  $Yb^{3+}$  does take place in  $CaSc_2O_4$  host lattice. However, more detailed measurements are required in order to understand the energy transfer mechanism clearly. The schematic energy level diagrams with the involved ET process are depicted in Figure 4, in which the proposed energy band diagram of  $Ce^{3+}$  is based on the PLE and PL spectra above, and that of  $Yb^{3+}$  is from the report of  $Yb^{3+}$  singly doped  $CaSc_2O_4$  sample by Gaume et al.<sup>26</sup> As shown in the scheme, the ET from the  $Ce^{3+}$  to the  $Yb^{3+}$  has two possible routes. (1) Cooperative ET, the energy of  $Ce^{3+}$  is



Figure 4. Schematic energy level diagrams of  $\mathrm{Ce}^{3+}$  and  $\mathrm{Yb}^{3+}$  with the involved ET process.

simultaneously transferred to two nearby Yb<sup>3+</sup> ions, resulting in two emitting photons of Yb<sup>3+</sup> with a wavelength around 1000 nm. (2) Single-step ET, it yields only a single NIR photon for each blue photon excitation. It should be noted that a secondorder ET is typically 10<sup>3</sup> times less possible than a first-order one, which makes cooperative ET occur with extremely low probability.<sup>28</sup> Recently, Meijerink et al. revealed that a singlestep ET instead of a cooperative process took place in the Ce<sup>3+</sup>/Yb<sup>3+</sup> codoped YAG material.<sup>6</sup>

For a cooperative ET process, the donor  $Ce^{3+}$  transfers its energy simultaneously to two Yb<sup>3+</sup> ions nearby. For the pair of acceptors, the energy transfer rate ( $W_{COOP}$ ) is proportional to the square of acceptor concentration.  $W_{COOP}$  can be simply written as  $W_{COOP} = Cx^2$ , where C represents the energy transfer coefficient and x is the Yb<sup>3+</sup> concentration. For single-step ET process, the Ce<sup>3+</sup> ion transfers its energy to only one Yb<sup>3+</sup> ion. Considering that the energy transfer rate ( $W_{S-S}$ ) is proportional to the acceptor concentration, it can be taken as  $W_{S-S} = C'x$ .

To judge the process of the ET from  $Ce^{3+}$  ions to  $Yb^{3+}$  ions in  $CaSc_2O_4$  sample, we analyzed the dependence of the energy transfer rate upon  $Yb^{3+}$  ion concentration. Figure 5 shows the decay curves of  $Ce^{3+}$ :  $Sd \rightarrow 4f$  emission at S11 nm in  $CaSc_2O_4$ :  $1\%Ce^{3+}/x\%Yb^{3+}$  (x = 0, 1, 2, 5, 10, 15) under 466 nm



**Figure 5.** Decay curves of Ce<sup>3+</sup>: 5d  $\rightarrow$  4f emission under 466 nm excitation in CaSc<sub>2</sub>O<sub>4</sub>: 1%Ce<sup>3+</sup>/x%Yb<sup>3+</sup> (x = 0, 1, 2, 5, 10, 15).

excitation. As Yb<sup>3+</sup> concentration increases, the decays rapidly speed up confirming remarkable Ce<sup>3+</sup>  $\rightarrow$  Yb<sup>3+</sup> ET. The lifetimes  $\tau_0$  and  $\tau$  of Ce<sup>3+</sup> in Ce<sup>3+</sup> singly and Ce<sup>3+</sup>/Yb<sup>3+</sup> doubly doped samples are calculated by integrating the area under the corresponding decay curves with the normalized initial

Table 1. Relationship of Ce<sup>3+</sup> Lifetime and Energy Transfer Efficiency ( $\eta_{\text{ETE}}$ ) with Different Concentrations of Yb<sup>3+</sup> Ions

Yb <sup>3+</sup> concentrations (%)	$\tau$ (ns)	$\eta_{\mathrm{ETE}}$ (%)
0	27.09	0
1	24.02	11
2	22.82	16
5	16.68	38
10	14.89	45
15	11.36	58

intensity. The lifetime values are shown in Table 1. The energy transfer rate  $(W_{Ce-Yb})$  can be obtained by

$$W_{\rm Ce-Yb} = \frac{1}{\tau} - \frac{1}{\tau_0} \tag{1}$$

In this way, we obtained the  $Yb^{3+}$  ion concentration dependence of the energy transfer rates and plotted in a double-logarithmic diagram, as shown in Figure 6. When the



**Figure 6.** Plot (log–log) of the energy transfer rate versus the Yb<sup>3+</sup> concentration in CaSc<sub>2</sub>O<sub>4</sub>: 1%Ce<sup>3+</sup>/x%Yb<sup>3+</sup> (x = 0, 1, 2, 5, 10, 15).

concentration of Yb<sup>3+</sup> ions increases from 1% to 15%, the slope is found to be around 1, indicating the ET from Ce<sup>3+</sup> ions to Yb<sup>3+</sup> ions is dominated by the single-step ET process rather than a cooperative ET process. On the basis of above data and analysis, we propose that its probability is low for Ce<sup>3+</sup>  $\rightarrow$  Yb<sup>3+</sup> ET to get a one-to-two photon process of DC in CaSc<sub>2</sub>O<sub>4</sub> material.

With Yb<sup>3+</sup> concentration increasing, Ce<sup>3+</sup> lifetime gradually decreases due to the ET from Ce<sup>3+</sup> to Yb<sup>3+</sup>, as shown in Figure 5. The variable distribution of Yb<sup>3+</sup> ions around the Ce<sup>3+</sup> ions leads to the variety of energy transfer rates; thus, all decay curves exhibit nonexponential characteristics. The energy transfer efficiency ( $\eta_{\text{ETE}}$ ) can be determined from decay curves. The  $\eta_{\text{ETE}}$  is defined as the ratio of Ce<sup>3+</sup> depopulated by ET to Yb<sup>3+</sup> over the total number of Ce<sup>3+</sup> excited. By dividing the integrated intensity of the decay curve of Ce<sup>3+</sup>/Yb<sup>3+</sup> codoped

 $\rm CaSc_2O_4$  sample by that of the  $\rm Ce^{3+}$  single doped one, the  $\eta_{\rm ETE}$  is obtained as a function of Yb^{3+} concentration

$$\eta_{\text{ETE},x\%\text{Yb}} = 1 - \frac{\tau}{\tau_0} \tag{2}$$

The estimated  $\eta_{\rm ETE}$  of samples presented in Table 1 exhibit a gradual increase with increasing Yb<sup>3+</sup> concentration. It is worth mentioning that  $\eta_{\rm ETE}$  reaches 58% for CaSc<sub>2</sub>O<sub>4</sub>: 1%Ce<sup>3+</sup>/15% Yb<sup>3+</sup> sample.

Figure 7 shows the normalized PLE and PL spectra of the doping concentration optimized  $CaSc_2O_4$ :  $1\%Ce^{3+}/5\%Yb^{3+}$ .



**Figure 7.** Normalized PLE spectrum of  $CaSc_2O_4$ :  $1\%Ce^{3+}/5\%Yb^{3+}$  (green line), PL spectra of  $CaSc_2O_4$ :  $1\%Ce^{3+}/5\%Yb^{3+}$  (red line), and YAG:  $1\%Ce^{3+}/5\%Yb^{3+}$  (black line) excited at 470 nm with the same power density. AM 1.5 G solar spectrum (blue line) and spectral response of c-Si (purple line) are the background.

The background is the AM 1.5 G solar spectrum and the spectral response of c-Si. As shown in Figure 7, the PLE spectrum of CaSc<sub>2</sub>O<sub>4</sub>: 1%Ce<sup>3+</sup>/5%Yb<sup>3+</sup> consists of a blue PLE band around 466 nm corresponding to the transition from the ground state to the lowest-lying 5d state, which agrees with the maximum photon flux region of the solar spectrum and one ultraviolet (UV) PLE band at 320 nm corresponding to the upper 5d states, respectively. Besides, CaSc<sub>2</sub>O<sub>4</sub>: 1%Ce<sup>3+</sup>/5% Yb<sup>3+</sup> gives an intense NIR emission at around 976 nm, matching well with the optimal spectral response of c-Si solar cells. Comparing with doping concentration also optimized typical phosphor YAG: 1%Ce<sup>3+</sup>/5%Yb<sup>3+</sup>, doping concentration optimized CaSc<sub>2</sub>O<sub>4</sub>: 1%Ce<sup>3+</sup>/5%Yb<sup>3+</sup> shows the stronger NIR emission. The NIR emission of CaSc<sub>2</sub>O<sub>4</sub> sample from 900 to 1100 nm is enhanced by a factor of 2.4 excited at 470 nm with the same power density, which corresponds to the best excitation of Ce<sup>3+</sup> in YAG: Ce<sup>3+</sup> sample.<sup>29</sup> Furthermore, the main emission peak of Yb<sup>3+</sup> in the  $CaSc_2O_4$  around 976 nm shows better agreement with the optimal spectral response of c-Si compared with that in YAG around 1028 nm. The NIR emission of Yb<sup>3+</sup> can be used more effectively in the CaSc<sub>2</sub>O<sub>4</sub>. Therefore,  $CaSc_2O_4$  is a promising solar spectral converter for achieving efficient NIR emission for c-Si solar cells.

# CONCLUSIONS

In summary, an efficient NIR DC has been demonstrated for the  $CaSc_2O_4$ :  $1\%Ce^{3+}/5\%Yb^{3+}$  sample. Compared with that in doping concentration optimized YAG:  $1\%Ce^{3+}/5\%Yb^{3+}$ , the NIR emission from 900 to 1100 nm is enhanced by a factor of

2.4. The main emission peak of Yb<sup>3+</sup> in the CaSc<sub>2</sub>O<sub>4</sub> around 976 nm matches well with the optimal spectral response of the c-Si solar cell. Furthermore, the mechanism of the DC process in CaSc<sub>2</sub>O<sub>4</sub>: Ce<sup>3+</sup>/Yb<sup>3+</sup> samples has been investigated carefully. On the basis of the analysis of the dependence of the energy transfer rate on Yb<sup>3+</sup> ion concentration, we propose that the ET from Ce<sup>3+</sup> ions to Yb<sup>3+</sup> ions occurs mainly by the single-step ET process. The estimated maximum energy transfer efficiency ( $\eta_{\text{ETE}}$ ) reaches 58% in the CaSc<sub>2</sub>O<sub>4</sub>: 1%Ce<sup>3+</sup>/15%Yb<sup>3+</sup> sample. The current experimental results indicate that CaSc<sub>2</sub>O<sub>4</sub>: Ce<sup>3+</sup>/Yb<sup>3+</sup> sample is a promising DC material for application in solar cells.

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# Notes

The authors declare no competing financial interest.

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